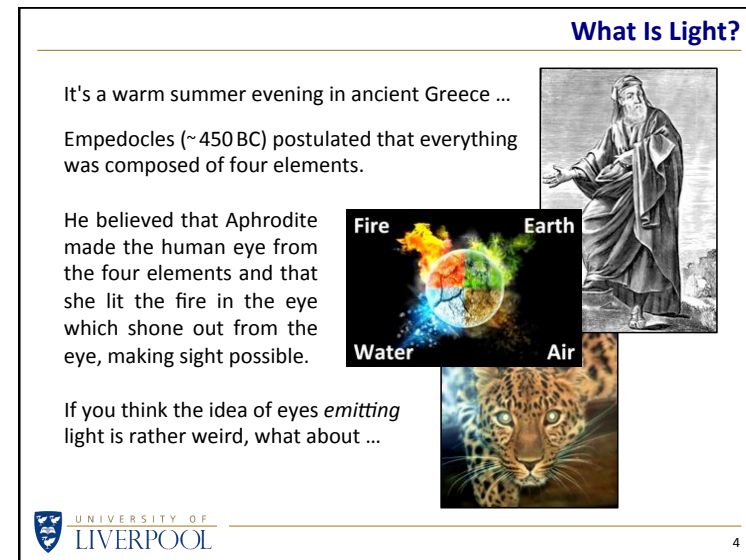
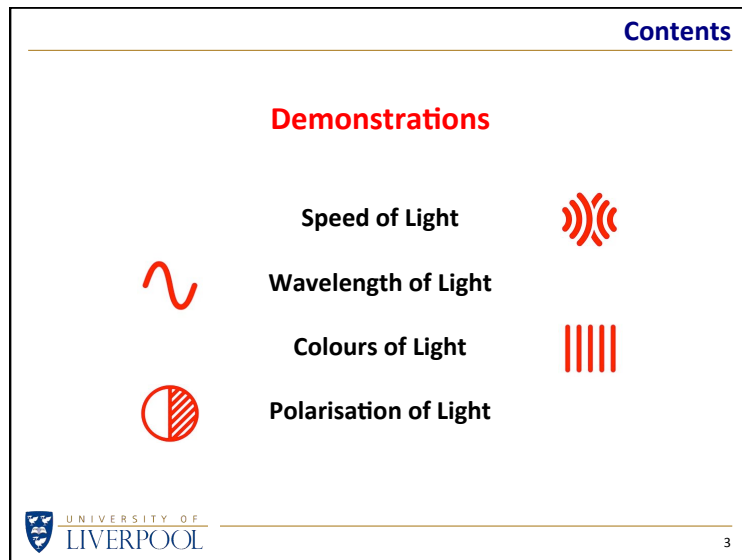
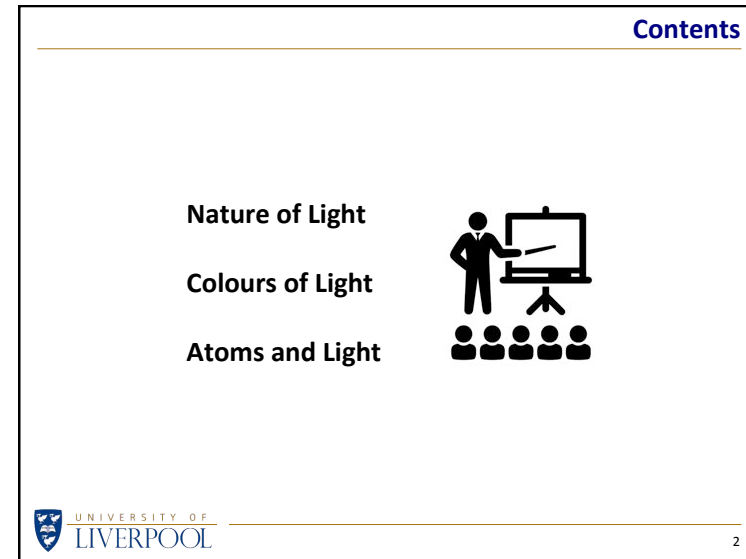
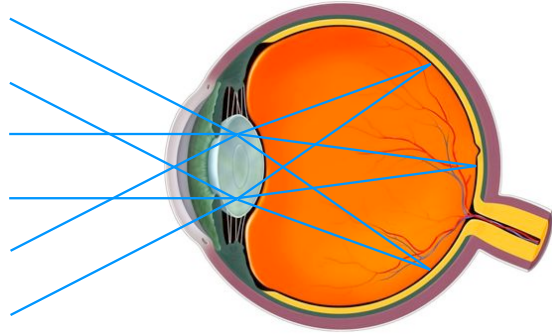


Fiat Lux



Fiat Lux

How Do We See?



Speed of Light

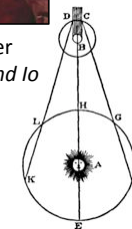
1638 Galileo Galilei
Lanterns on hilltops



1728 James Bradley
Stellar aberration



1676 Ole Rømer
Jupiter and Io



$c = 300,000 \text{ km/s}$

Isaac Newton



Newton's work on optics in the 1660s established that white light is composed of many colours.

Arguments raged as to the nature of light – should it be considered as a stream of particles ('corpuscular') or as a wave?

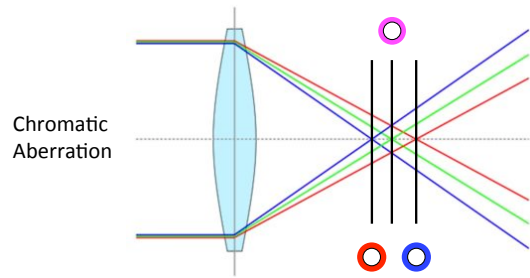
(Curiously, modern physics developed centuries later gives the answer – both!)

Making Rainbows



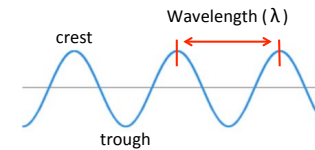
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Unwanted Rainbows

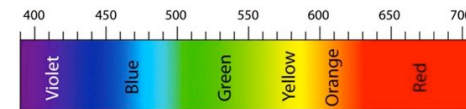


Light as a Wave

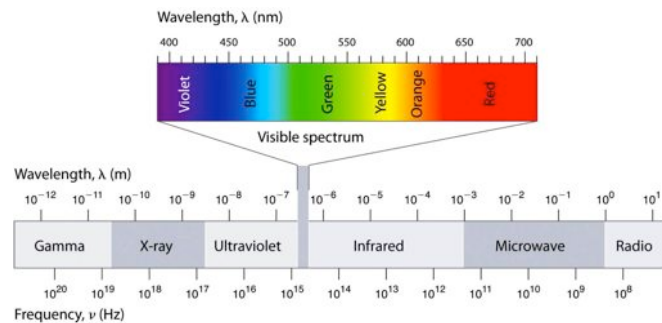
When light passes through an aperture, such as an optical instrument like a telescope or a narrow slit, it reveals its wave nature.



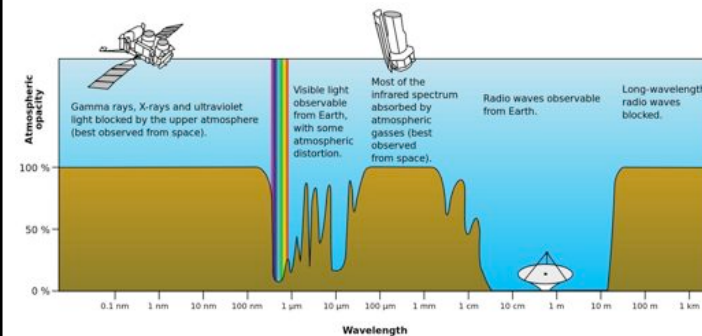
Wavelength (nm)



Electromagnetic Spectrum

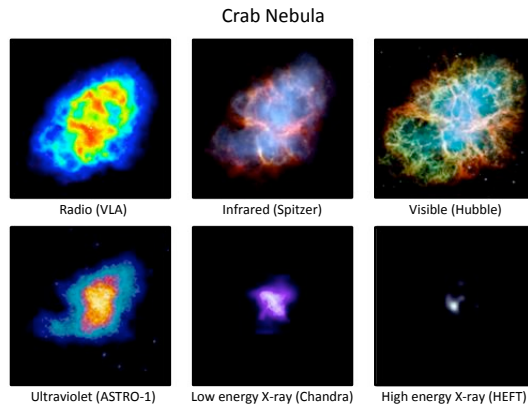


Electromagnetic Spectrum

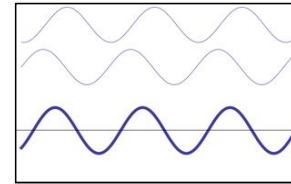


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Electromagnetic Spectrum



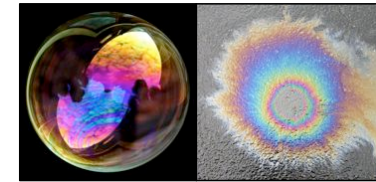
Wave Interference



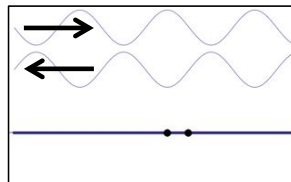
Animation courtesy of Dr Dan Russell, Pennsylvania State University

Just like waves on the surface of water, if two light waves meet and their crests and troughs fall on top of each other then they reinforce each other. However, if the crest of one wave meets the trough from another then they cancel each other out. We call this **interference**.

You will have seen some of the results of interference in soap bubbles or oil spills.



Standing Waves



Animation courtesy of Dr Dan Russell, Pennsylvania State University

Waves that travel in one direction and meet similar waves travelling in the opposite direction will produce **standing waves**.

Some points ('hot spots' like the right dot) see big variations in the waves. Other points ('dead spots' like the left dot) see no waves at all.

Hence if we can create standing waves inside an object, some regions will feel the wave and some won't. Sending microwaves (radio-frequency light waves) into food will result in 'hot spots' being cooked and 'dead spots' remaining uncooked.

Demonstration



Speed of Light

1. Take an ordinary microwave oven

2. Remove the rotating turntable

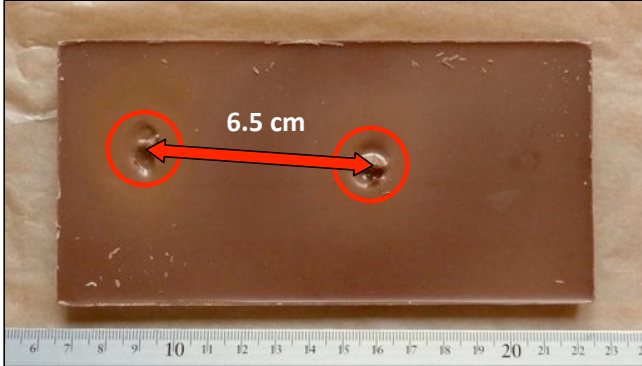
3. Put a chocolate bar into the oven (it also works with cheese but it is (a) not as clear and (b) a waste of cheese)

4. Cook for 15–20 seconds

5. Measure the distance between the melted regions



 **Demonstration**



6.5 cm

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 **Demonstration**

Speed of Light

6. The wavelength of the microwaves = 2 x this distance = 13 cm

7. Note the frequency of the microwaves (on the back of the oven)

8. Speed = frequency x wavelength (*wiggles per second* x *length of each wiggle*)

9. Put in the numbers ...

Speed = 2450 MHz x 13 cm = **320,000** km/s

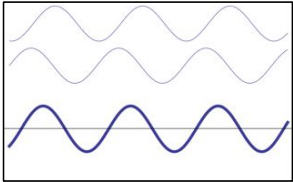
Actual value = **300,000** km/s



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Interference and Diffraction



Just like waves on the surface of water, if two light waves meet and their crests and troughs fall on top of each other then they reinforce each other. However, if the crest of one wave meets the trough from another then they cancel each other out. We call this **interference**.

Animation courtesy of Dr Dan Russell, Pennsylvania State University

When lots of waves interfere with each other we call this **diffraction**.

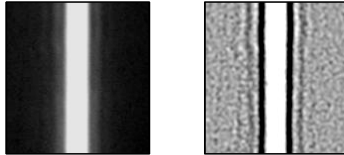
We don't usually notice diffraction effects unless we look very closely (for instance, with high magnification eyepieces in a telescope) or the light passes through very small apertures such as a slit of width of only a fraction of a millimetre.

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Diffraction

A wide slit produces diffraction 'fringes' that are closely spaced and a narrow slit produces wider (and more noticeable) diffraction fringes. This is why large telescope lenses or mirrors produce 'sharper' images.



↑↑↑ ↑↑↑

Enhancing the contrast makes the diffraction fringes easier to see*

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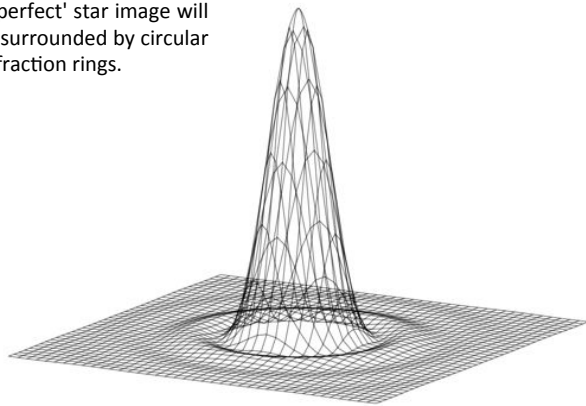
* www.liv.ac.uk/~sdb/ProcSyzSoc/Proc-Syz-Soc-2.pdf

20

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Diffraction

A 'perfect' star image will be surrounded by circular diffraction rings.



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Demonstration

Wavelength of Light

- Shine a bright light through a narrow slit onto a screen
- Observe diffraction fringes either side of the slit

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Demonstration

Wavelength of Light

Measure the width of central bright region

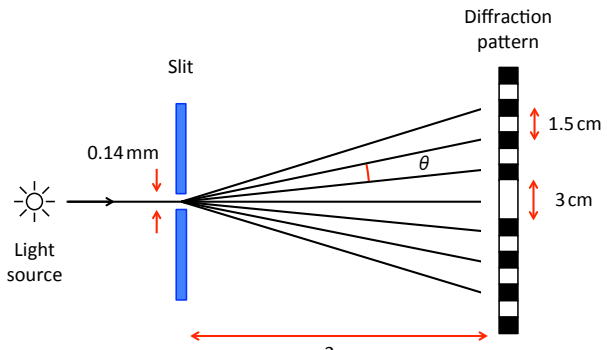


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Demonstration

Wavelength of Light



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Demonstration

Wavelength of Light

The angular separation of the fringes ($\theta \approx 1.5/300$) is given by the ratio of the wavelength of light to the slit width.

The angular separation of the fringes = $\frac{1.5}{300} = \frac{\text{wavelength}}{\text{slit width}} = \frac{\lambda}{0.14 \text{ mm}}$

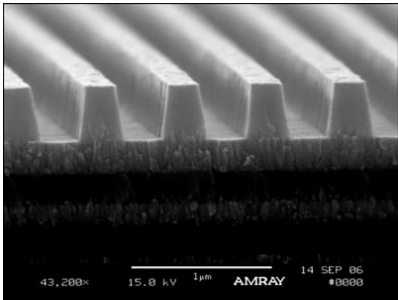
so the wavelength of light = $\frac{1.5}{300} \times 0.14 \text{ mm} = \mathbf{700 \text{ nm}}$

It ought to be about 650 nm, as the wavelengths of visible light span the range 400–700 nm (from blue to red) and a red laser was used in this demonstration.


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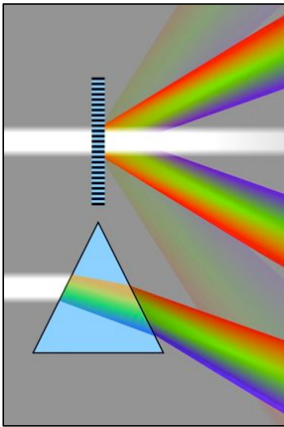
Diffraction Grating

A diffraction grating, usually a piece of glass etched with closely-spaced parallel lines, uses the interference of light waves to make a spectrum.




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Diffraction Grating



Some diffraction gratings are designed to make spectra by passing light through them, others are designed to make spectra by reflecting light.

Note that the colours with *smaller* wavelengths (at the blue end of the spectrum) are sent through *smaller* angles. This is the opposite to the way a prism works.


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Diffraction Grating

You have probably already seen a diffraction grating.

CDs and DVDs were not created because of the way they diffract light, but the closely-spaced tracks that are used to record the music or video data do diffract light just like a diffraction grating.



(This effect can be exploited by making a 'CD spectroscope' from a CD and an empty cereal box. *)

You have probably also seen diffraction effects in nature without realising it ...




* www.liv.ac.uk/~sdb/ProcSyzSoc/Proc-Syz-Soc-3.pdf
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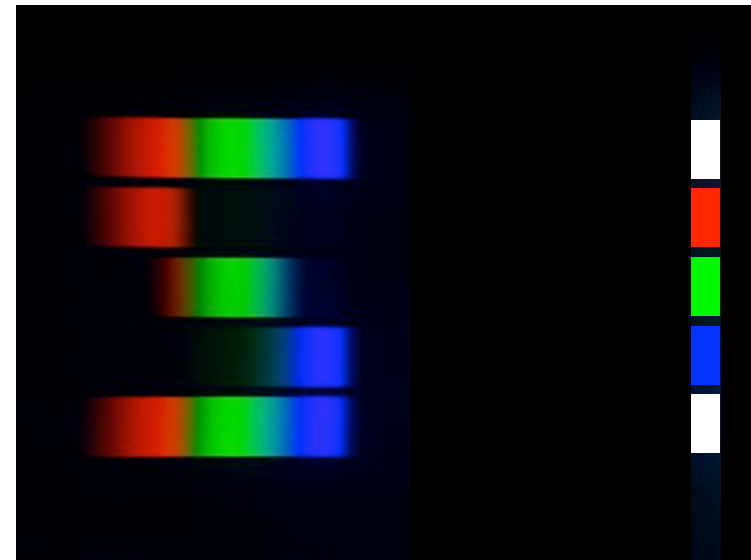


Demonstration

Spectrum of Light

- Make a spectrum using a prism
- Make a spectrum using a diffraction grating
- Look at different light sources using a diffraction grating

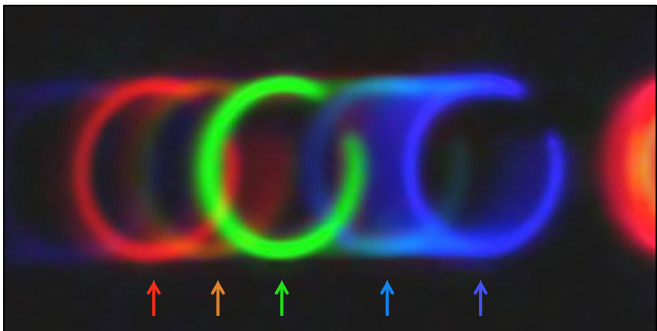
29





Demonstration

Fluorescent Light Bulb



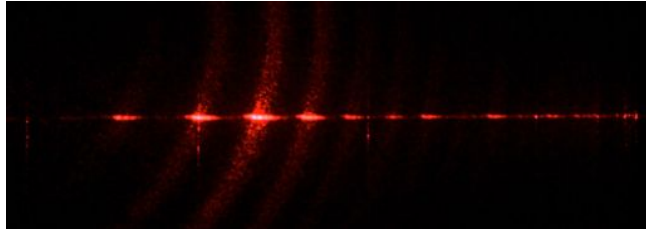
Eu Eu Hg Tb Hg

31



Demonstration

Laser + 6" Ruler

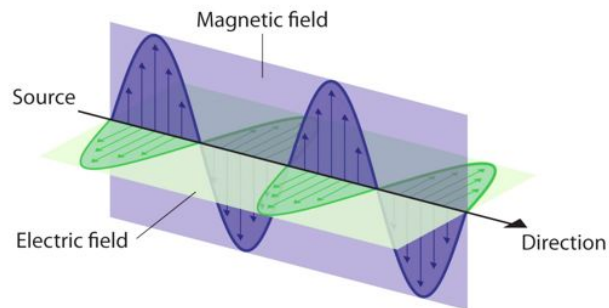


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Electromagnetic Wave

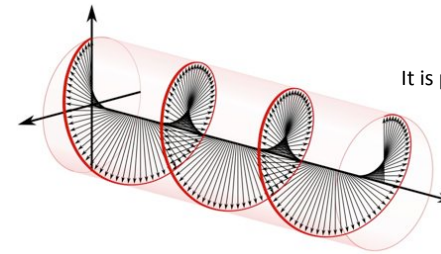
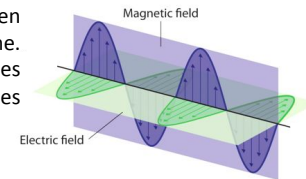
In 1865 James Clerk Maxwell showed that light is an electromagnetic wave. Electric fields and magnetic fields interact with each other and are interwoven to make a wave that can travel without a medium.



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Polarisation

The electric and magnetic fields (green and blue arrows) stay in the same plane. For instance, the magnetic field varies 'up-down' and the electric field varies 'left-right'. (*Linear Polarisation*)



It is possible to create light waves that corkscrew around the direction in which they travel.

(*Circular Polarisation*)

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Demonstration

Polarisation of Light

- Pass light through linear polarising filters (aka Polaroid™)
- Pass light through a calcite crystal
- Pass light through circular polarising filters (aka 3D TV glasses)

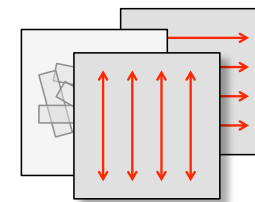
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Demonstration

Crossed Polarising Filters

Place two polarising filters together in perpendicular orientations so that no light is transmitted through them



Place a sheet of plastic between them and see what happens. If the plastic has strips of sticky tape on it, they appear coloured dependent on the thicknesses of the tape.

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**Demonstration**

Crossed Polarising Filters



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**Demonstration**

Calcite Crystal Double Refraction

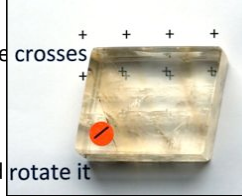
Place the calcite crystal over this text or the crosses

You see two images of everything

Place a polarising filter over the crystal and rotate it

You see that each image has a different polarisation

This is because light takes two different paths through the crystal depending on its polarisation



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**Demonstration**

Vampire Zebras



Zebras on the screen (left) have no reflection in the window (right)

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**Demonstration**

Left-Handed Beetle



Left - Circular

Right - Circular

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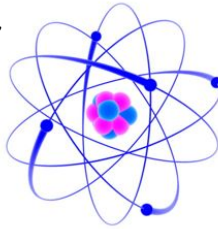
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Atoms and Light

Electromagnetic waves are created whenever charged particles are moved around.

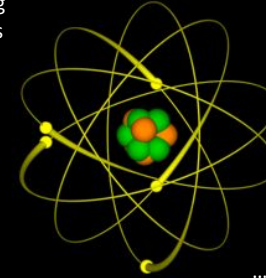
This can happen in a number of different ways, such as accelerating electrons or protons in a particle accelerator, but the smallest objects that involve moving charges are **atoms**.

Hence to understand how light is emitted or absorbed we should look at how atoms work.



Electrons in Atoms

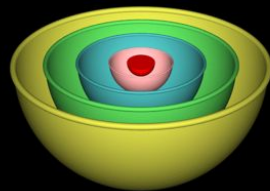
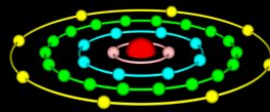
This might be how we imagine atoms with electrons buzzing around a nucleus like bees ...



... but it doesn't show us that all the electrons have different energies

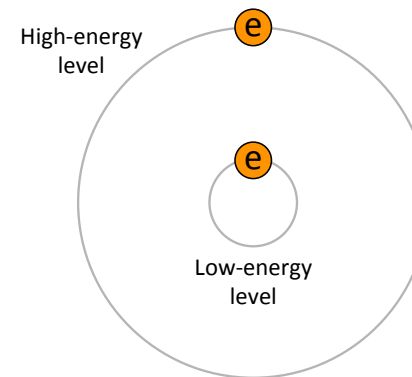
Electrons in Shells

It is better to think of the electrons in different sized orbits ...



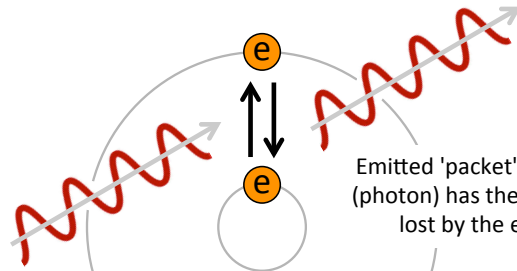
... or concentric shells surrounding the nucleus

Quantum Jumps



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Absorption and Emission



The electron can only absorb the energy of the photon if the energies are an EXACT match

Emitted 'packet' of light (photon) has the energy lost by the electron

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Photon Energy and Colour

According to Quantum Mechanics – the laws of Physics that describe how things work on very small scales – a photon of light with a given energy has a particular wavelength, and hence a particular colour.

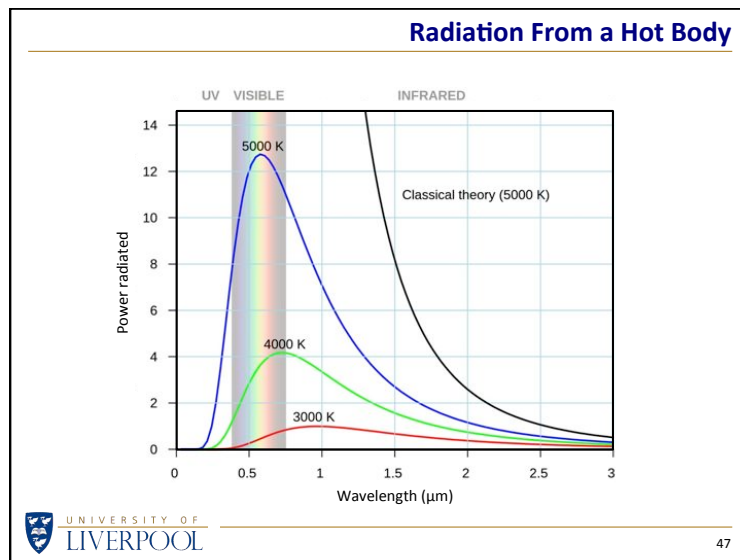
Lower ————— Energy —————> Higher



700nm <----- Wavelength -----> 400nm

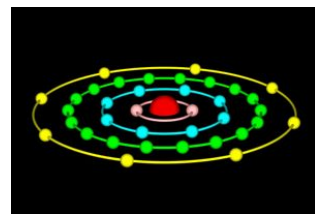
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Emission Spectrum

Every atom has many energy levels and so an electron can make quantum jumps with a number of specific energies that are characteristic of the atom.



Hence atoms emit or absorb photons with characteristic patterns of energies and hence colour spectra.

Emission



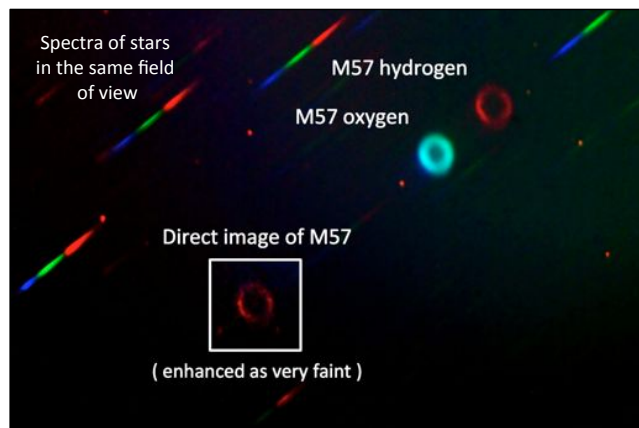
Absorption



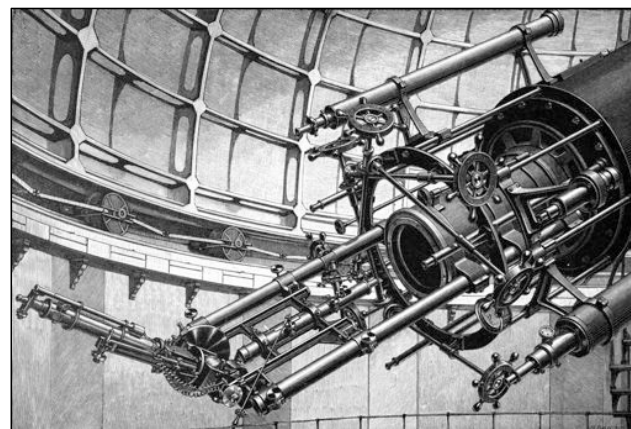
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Spectrum of M57 Ring Nebula



Spectroscope



Star Analyser 100

PHL PATON HAWKSLEY EDUCATION LTD
Manufacturer of diffraction gratings, spectroscopes and educational science

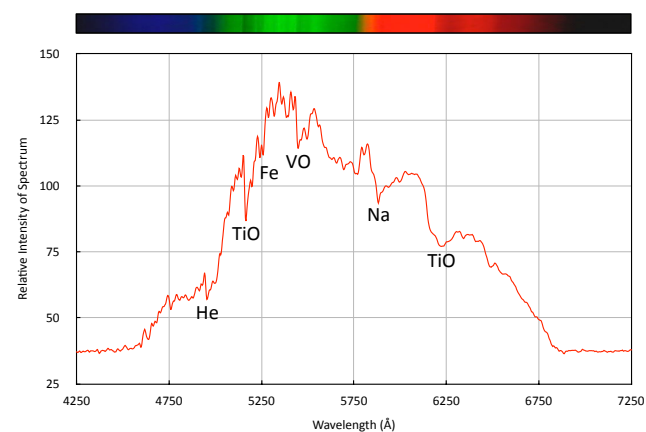
About Us Products Contact Us

The Star Analyser

With the Star Analyser... Turn this... Into this!

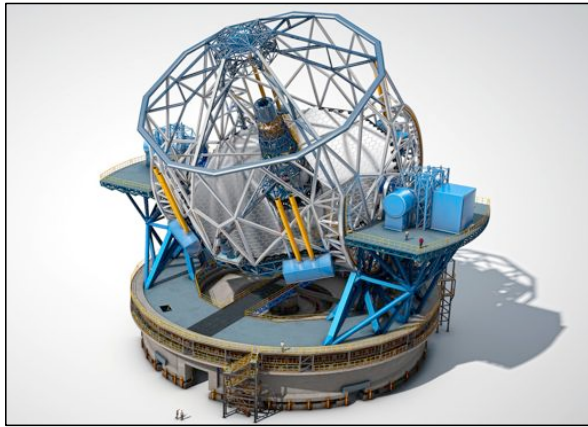
The STAR ANALYSER is used here with an 80mm reflector and an unmodified Philips Toucan Pro webcam, to reveal the telltale signature of molecules in the atmosphere of the magnitude +3.4 red giant Delta Virginis (spectral class M3 iii) (Graph plotted from spectrum image using Visual Spec software)

Spectrum of Betelgeuse



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European Extremely Large Telescope



Evolution of Telescopes

Telescopes have evolved with ever larger mirrors to collect the light.

Larger mirrors means better resolution, so **SHARPER** images.

Larger mirrors mean more light, so **BRIGHTER** images.

But remember that one of the most important instruments on these telescopes is the spectrograph, as without it we would not be able to understand what the light from these distant objects is telling us.

(For instance, note that the EELT has **3** cameras, but **7** spectrographs!)

Astronomy has come a long way since Auguste Comte (1798 – 1857):

"We will never know ... the composition of the stars"

Summary

You now have a better understanding of what light is.

You have seen demonstrations that have shown that ...

The speed of light is 300,000 km/s



The wavelength of light is 400–700 nm



Light can be split into its component colours



Light waves can be polarised in different ways

Now take a diffraction grating and go and experiment for yourself ...

Many thanks to roadies Sue and Rob
for helping with the demonstrations

fiat Lux

<http://www.liv.ac.uk/~sdb/Talks>

Dr Steve Barrett

BASoc 20 Feb 2017